

Cross-scale Observational Signatures of Magnetic Reconnection

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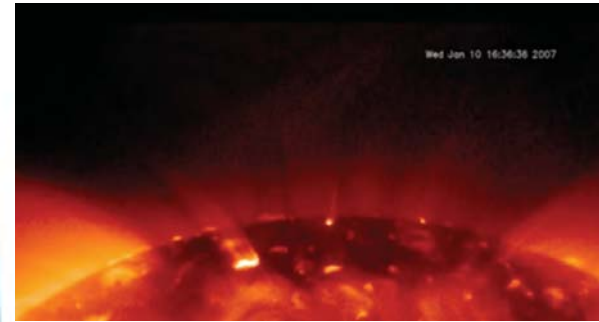
Variety of Signatures

Solar, Non-laboratory
...to name just a few

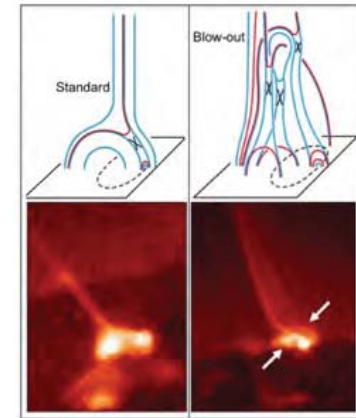


Example observations of reconnection

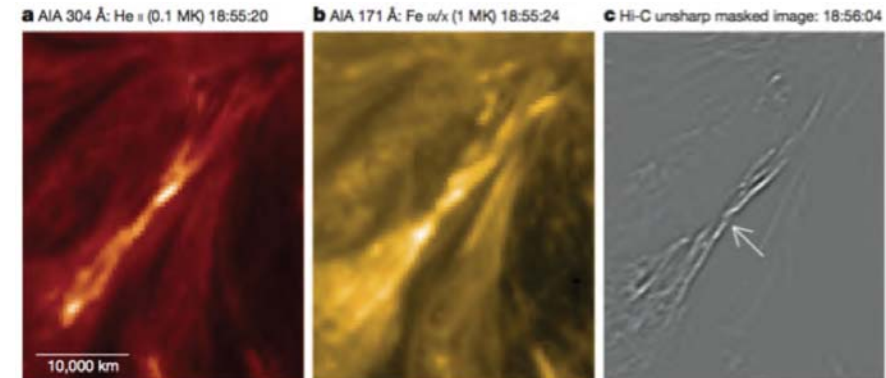
- ♦ Polar regions teeming with jets
- ♦ Field line unbraiding suggested by the Hi-C sounding rocket observations (193 Å)
- ♦ Field line spreading
- ♦ Flare initiation, etc.



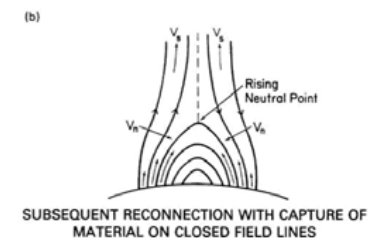
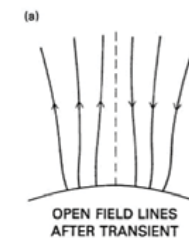
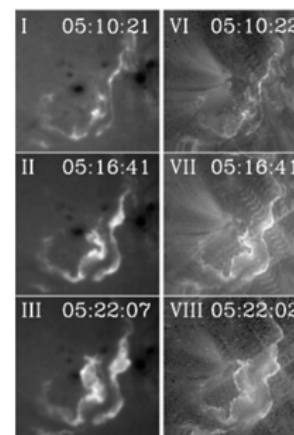
Cirtain et al 2007



Moore et al 2011;
Shimojo et al 1996



Cirtain et al 2013



Kopp & Pneuman 1976

Asai et al 2003

Current Sheet Reconnection

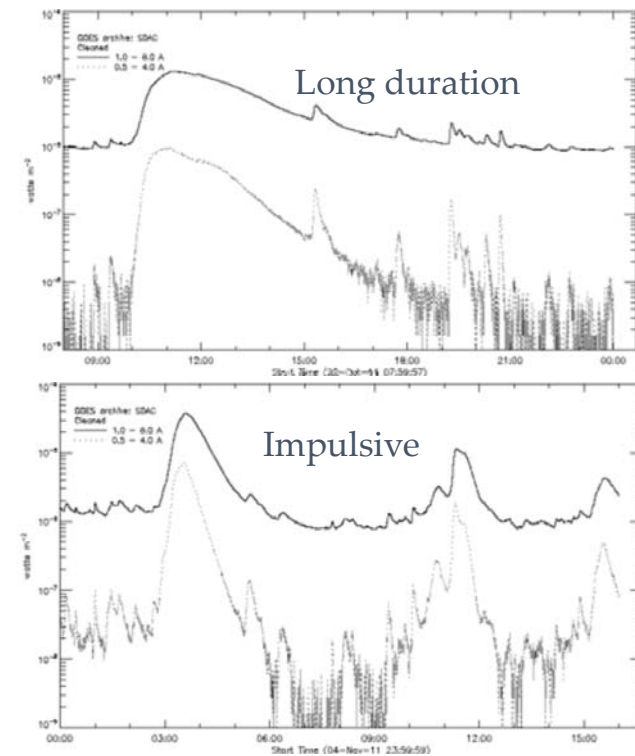
...during long duration solar flaring events



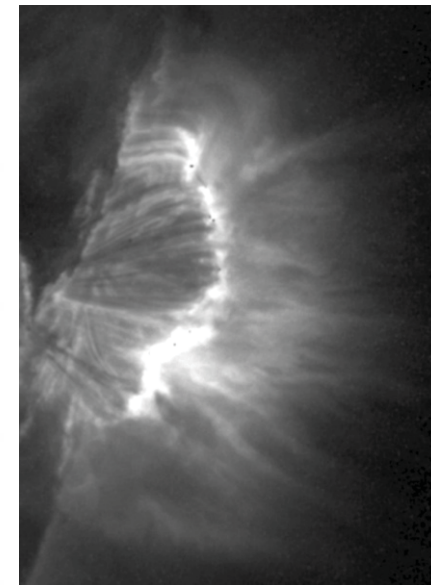
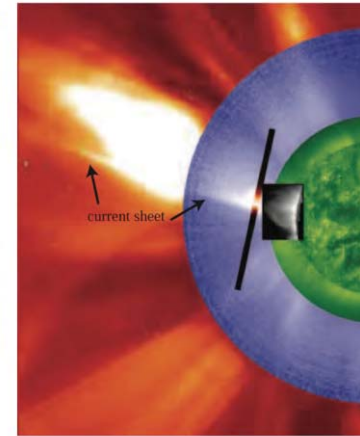
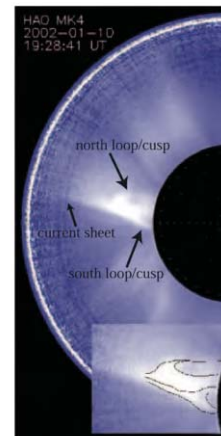
Long Duration Events

- Energy released for many hours
- CMEs
- Development of current sheets & supra-arcade fans

Example GOES Lightcurves



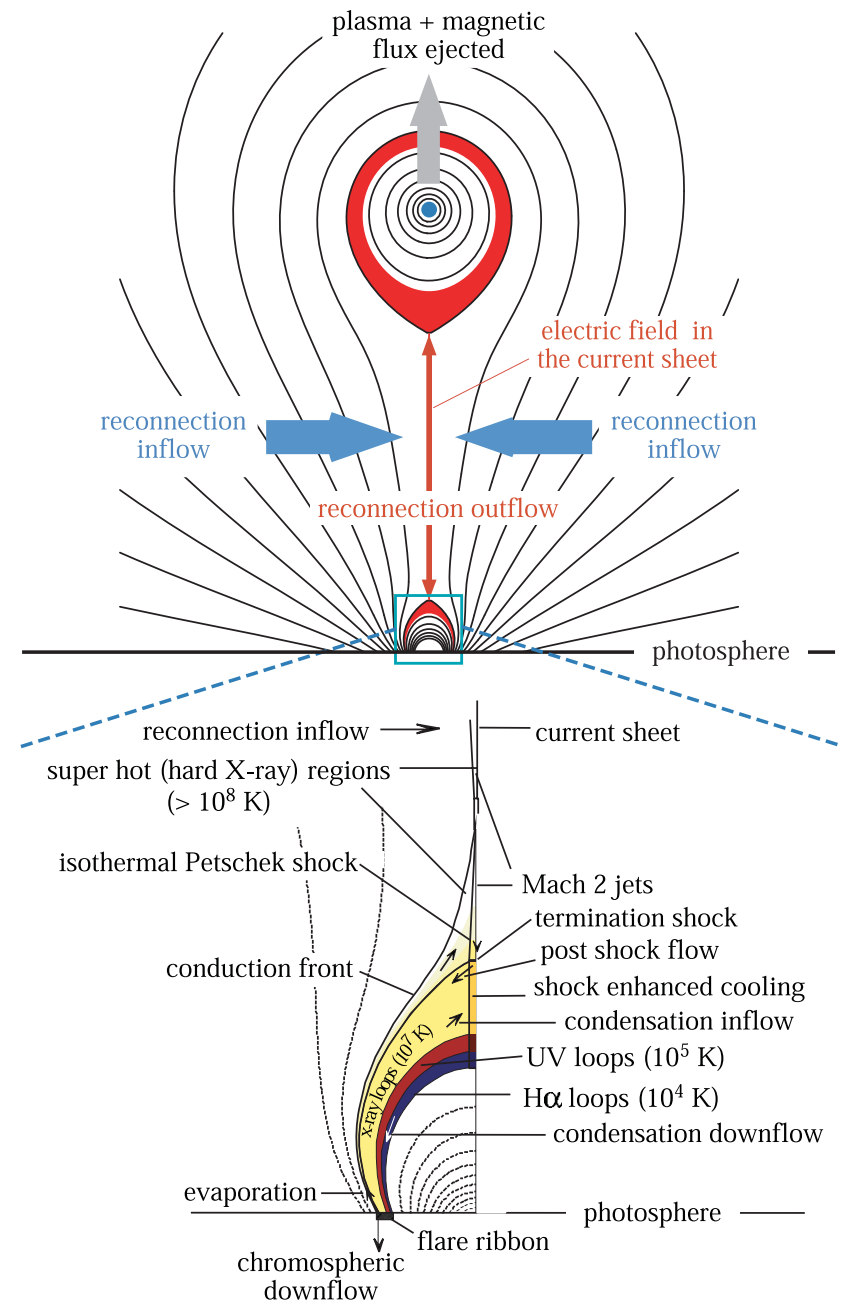
Ko et al 2003

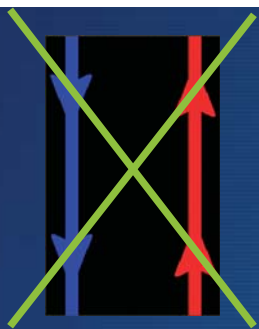


Savage & McKenzie 2011

Theory

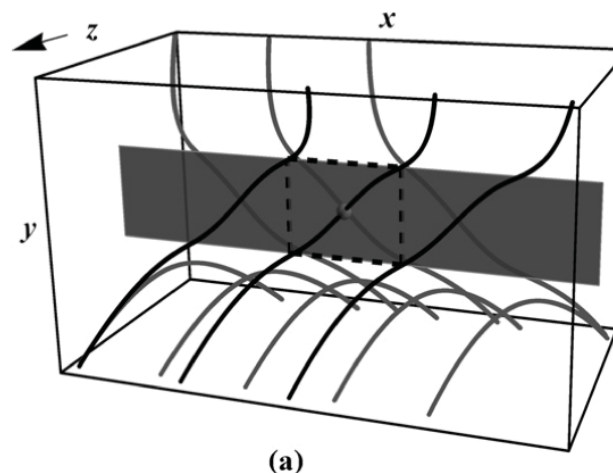
- Basic 2D ReX
- Magnetic Islands
 - Enablers or Comparable Outputs?
- CSHKP model, updated





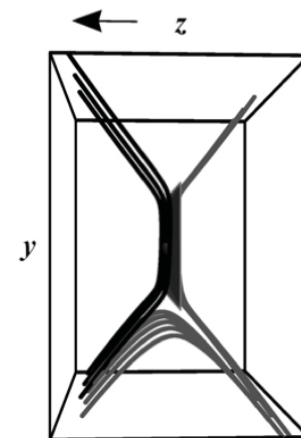
Theory

- 💧 Basic 2D ReX
- 💧 Magnetic Islands
 - 💧 Enablers or Comparable Outputs?
- 💧 CSHKP model, updated
- 💧 3D ReX for fast reconnection and plasmoids

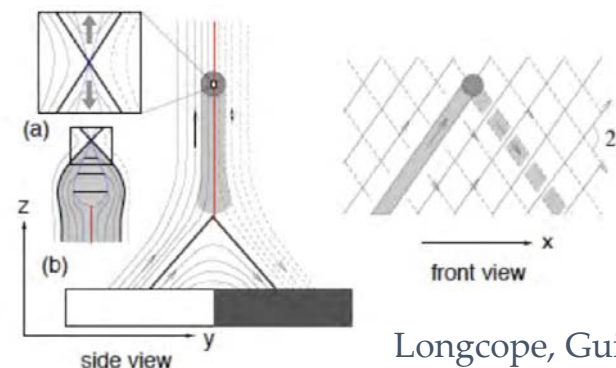


(a)

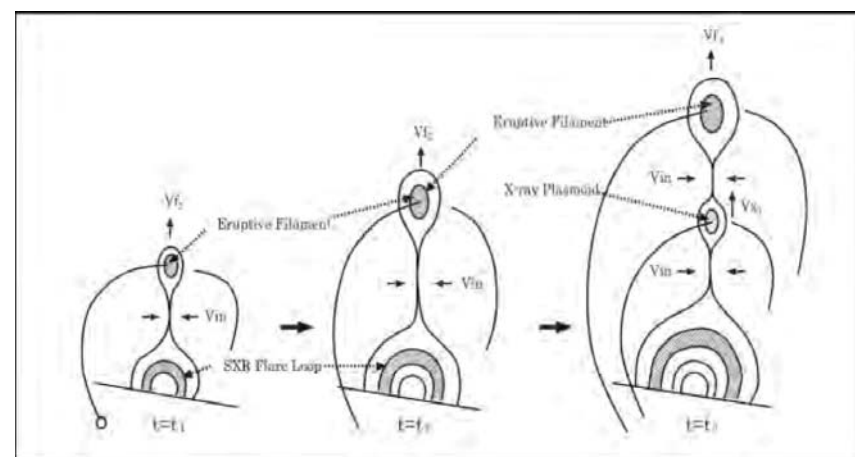
Guidoni & Longcope 2010



(b)



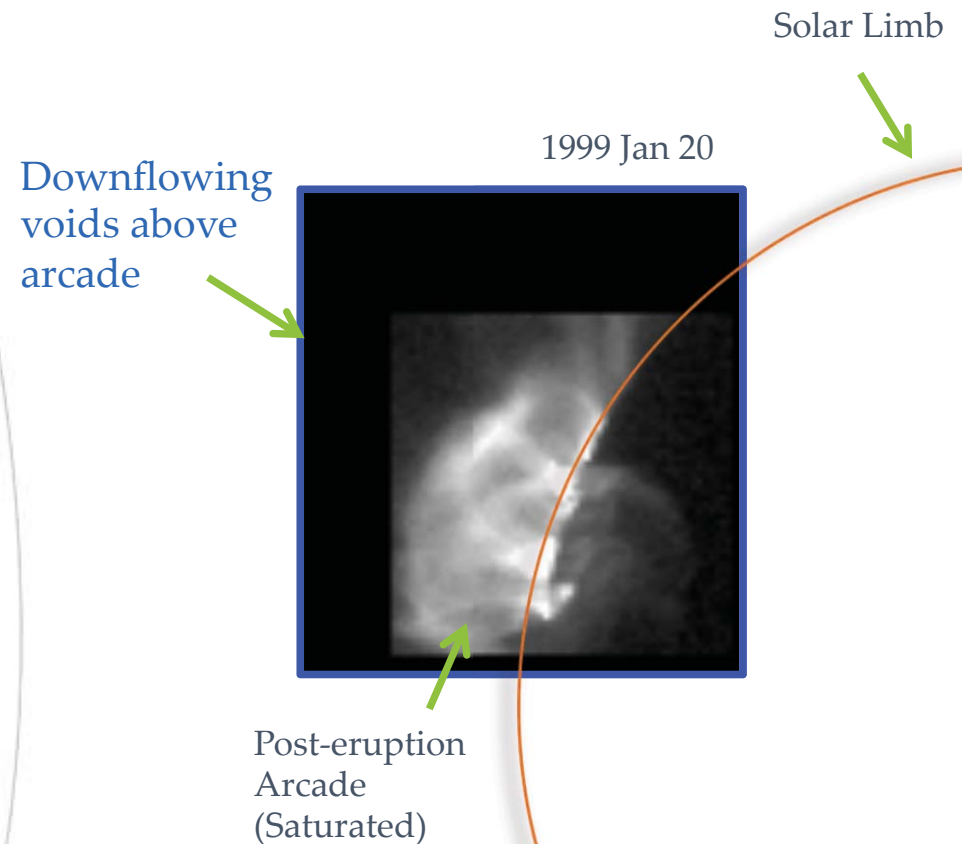
Longcope, Guidoni, & Linton 2009



Ohyaama & Shibata 2008

SXT

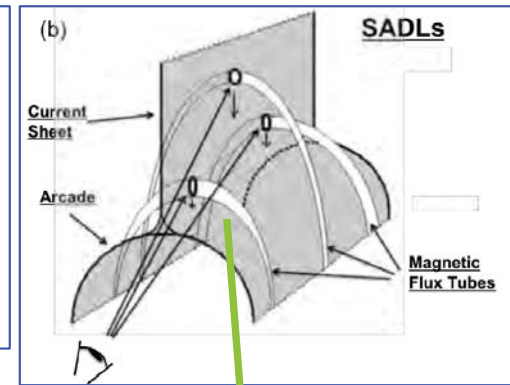
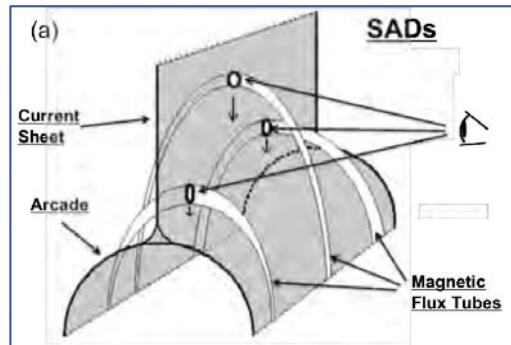
- First observation of downflowing voids above developing post-eruption arcades



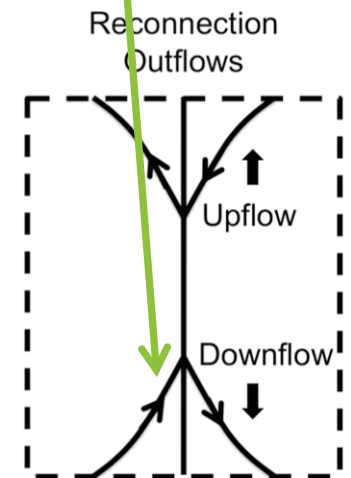
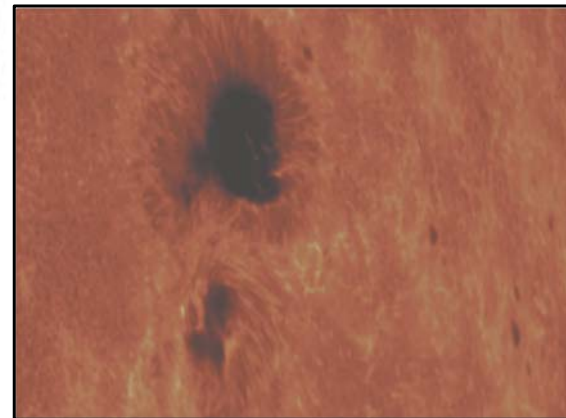
McKenzie & Hudson 1999

SADs / SADLs

- Teardrop shaped voids flowing sunwards through the voids
- Cross-sections of retracting flux tubes
 - reconnection outflows

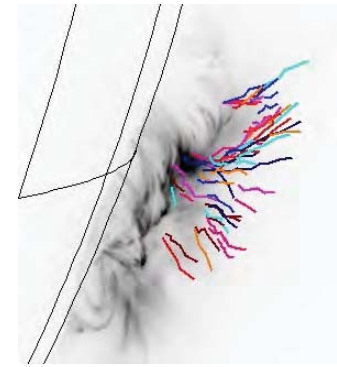
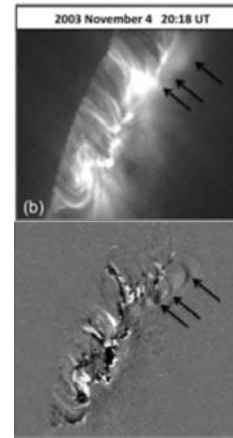
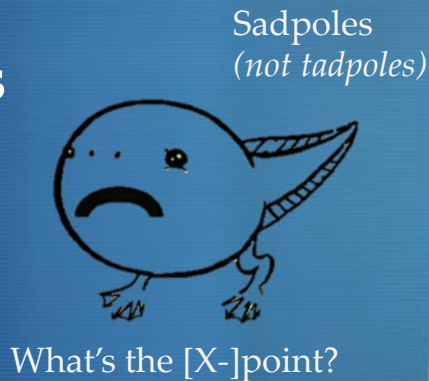


McKenzie & Hudson 1999
McKenzie & Savage 2009



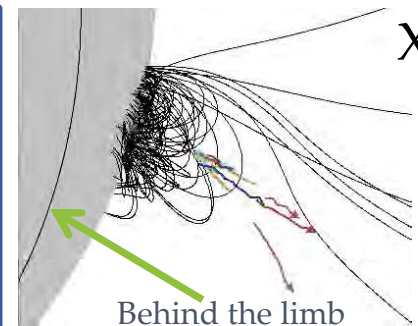
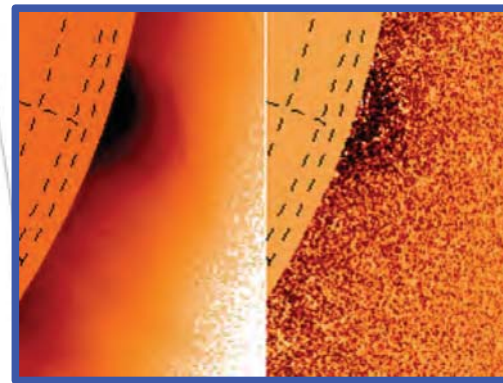
Observed across the spectrum

- 💧 X-class flares
- 💧 Voids
- 💧 Loops
- 💧 Disconnection Event
- 💧 “Plasmoids” & outflows
- 💧 Many hours after initiation in LASCO



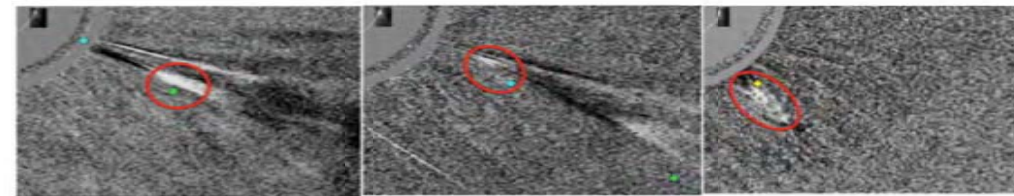
McKenzie & Savage 2009

TRACE

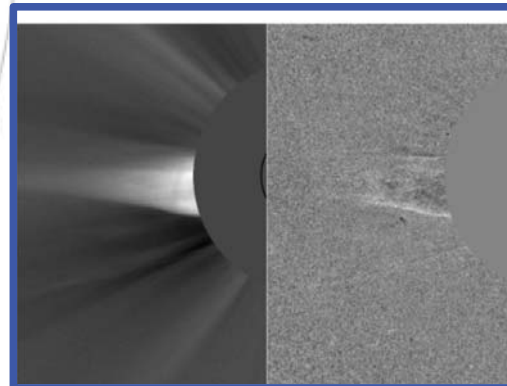


Savage et al. 2010

XRT



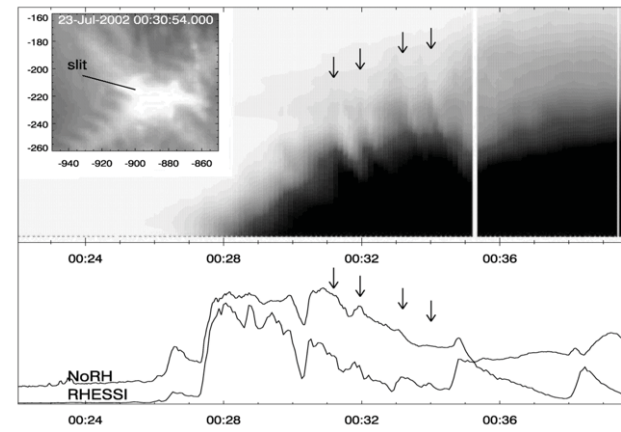
LASCO



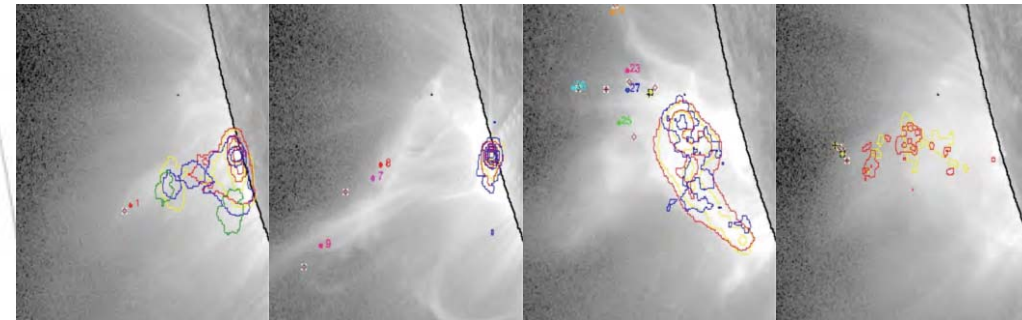
Observed across the spectrum

- Particle acceleration to footpoints (HXR's)
- Hot thermal plasma temporally and spatially associated with downflows
- Radio termination shocks
 - Nonthermal & coherent

RHESSI

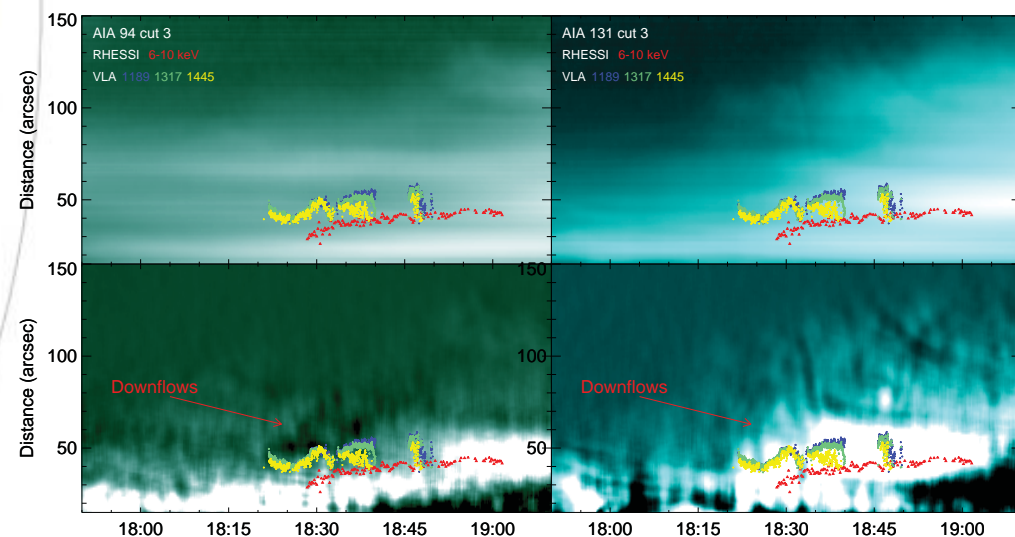


Asai et al 2004



Savage – 2010 Nov 3 flare

VLA



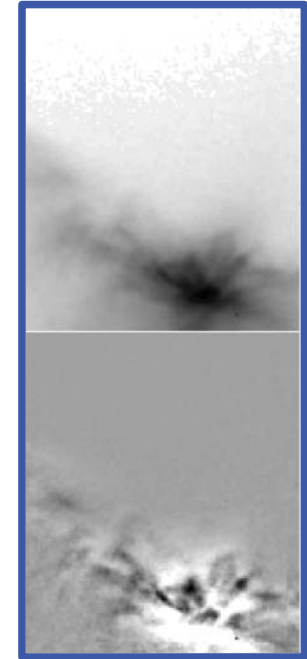
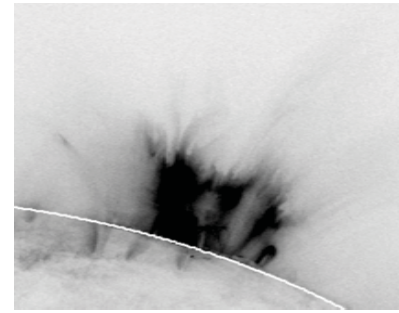
Chen et al. 2014 – 2012 Mar 3 flare, courtesy of T. Bastian

AIA

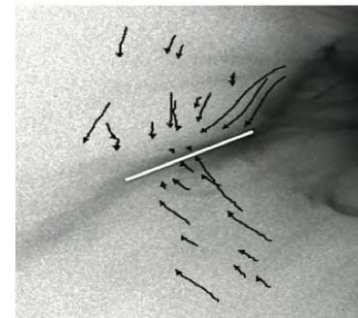
- Inflows & Outflows
- SADs & SADLs

2011 May 9

AIA



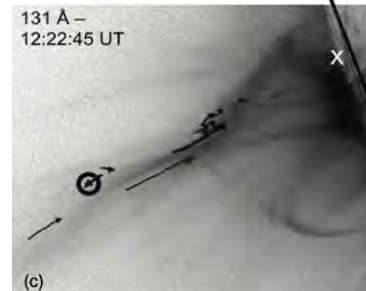
2010 Nov 3



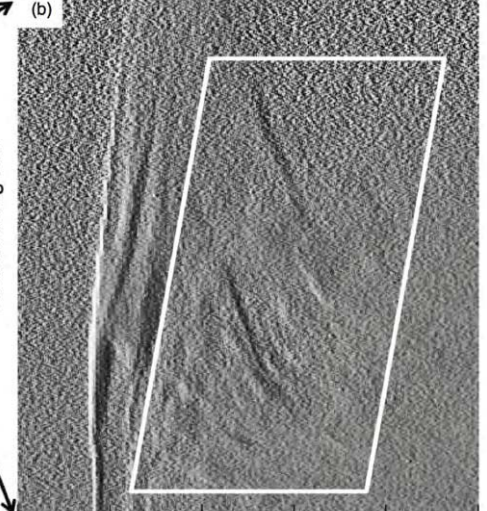
131 Å –
12:18:09 UT



131 Å –
12:22:45 UT



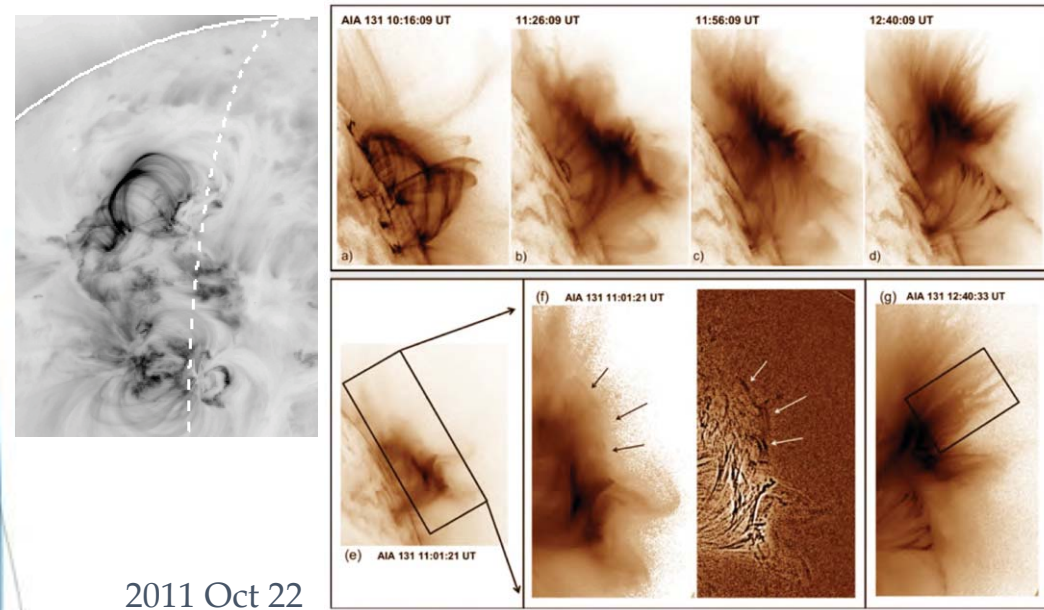
Position along Slit



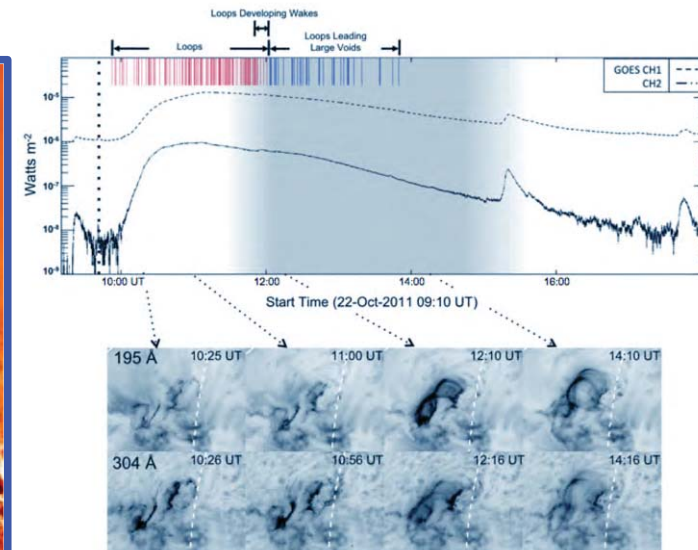
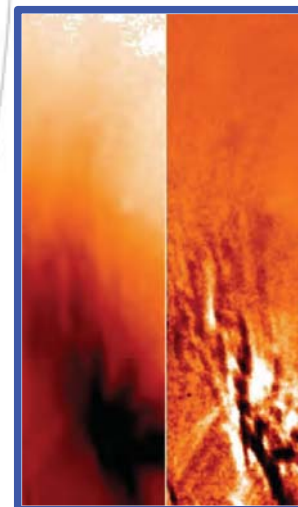
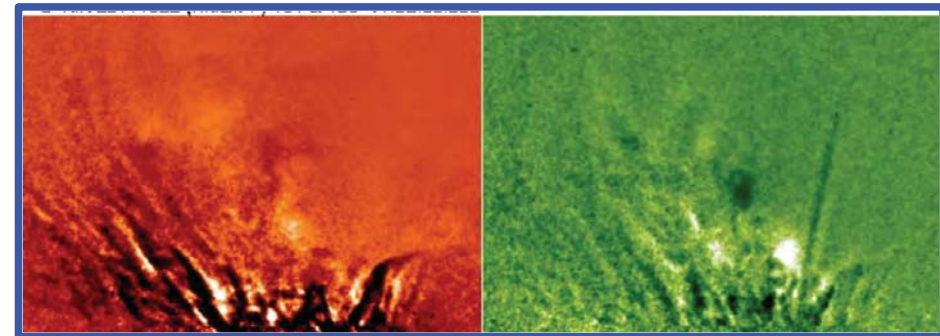
12:10:21 12:15:21 12:20:21 12:25:09 12:30:09 12:34:57
Time [s]

AIA

- Paradigm shift
 - Flare defying geometrical argument
 - Voids in current sheet behind thin retracting flux tubes vs. Cross-sections



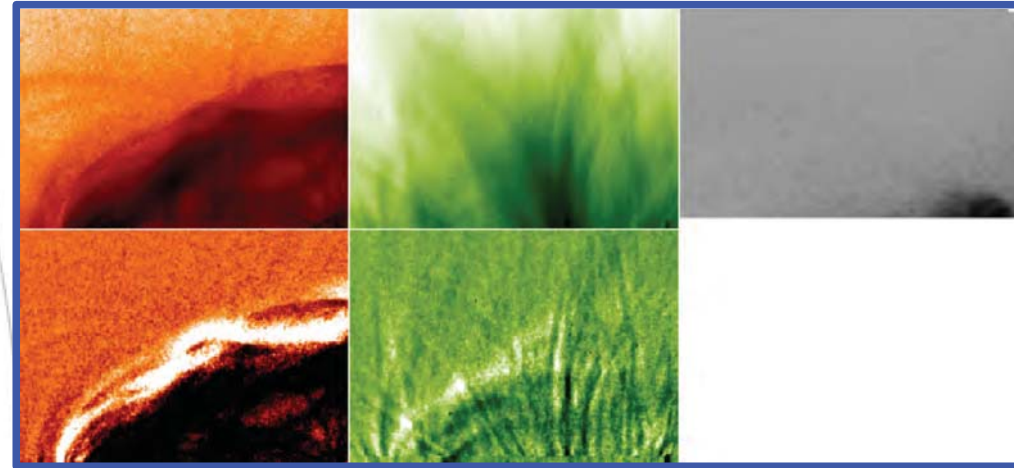
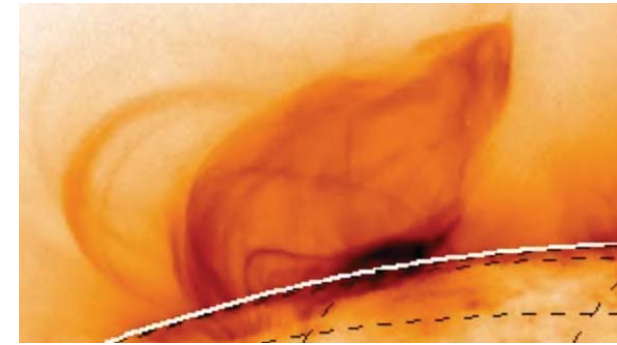
2011 Oct 22



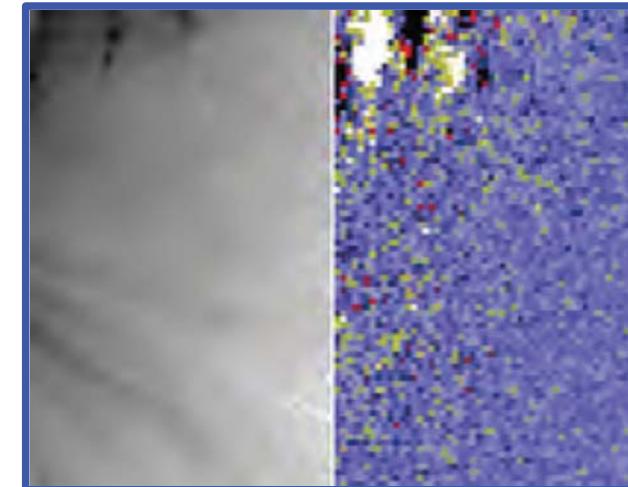
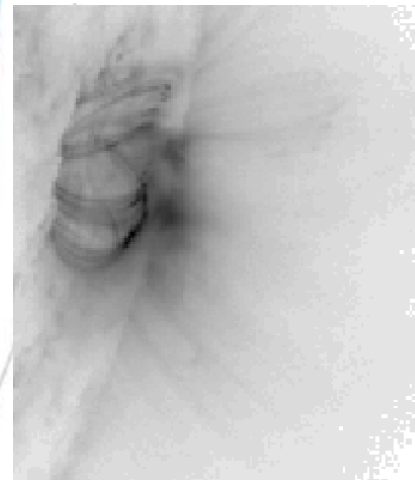
AIA

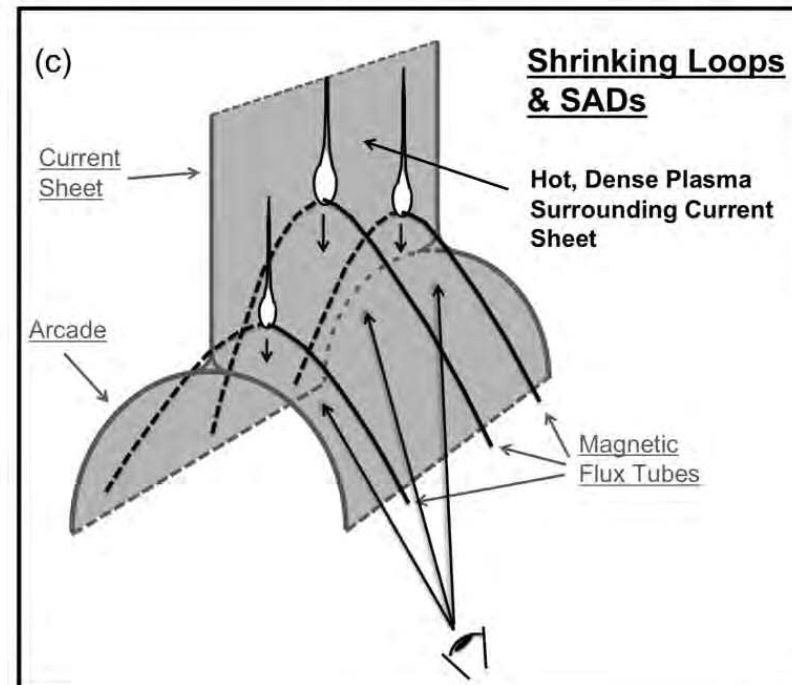
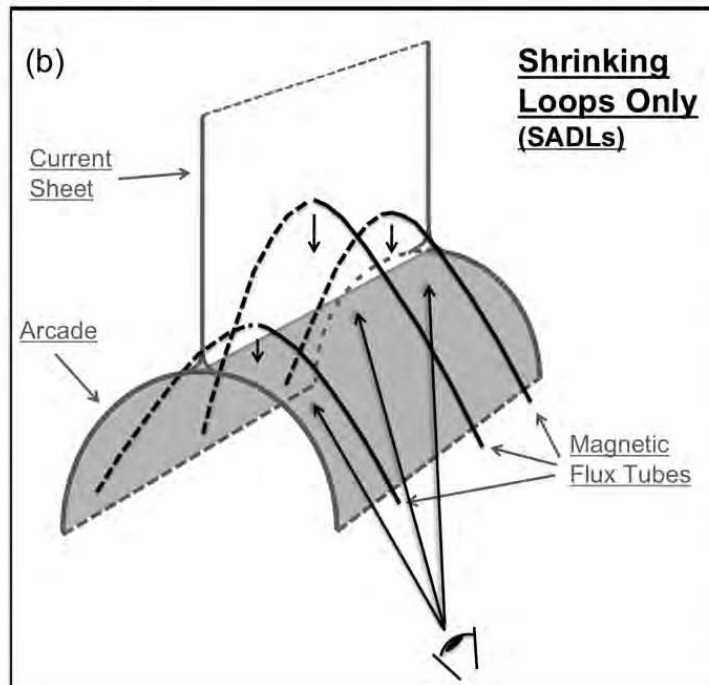
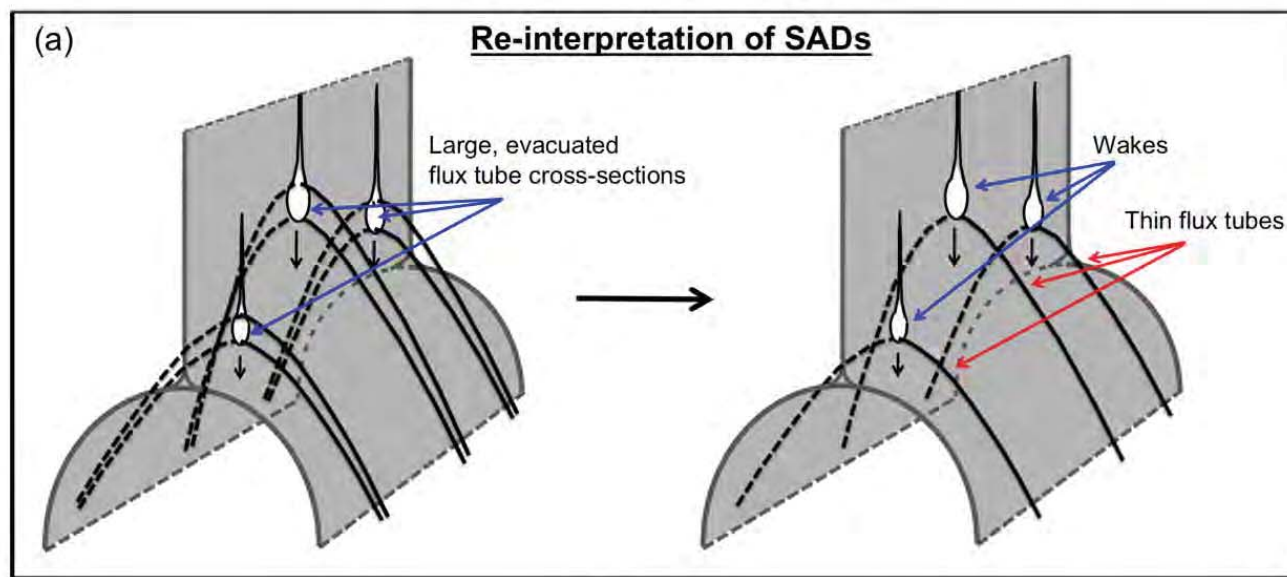
- Paradigm shift
 - Flare defying geometrical argument
 - Voids in current sheet behind thin retracting flux tubes vs. Cross-sections
- Mounting observational evidence

2012 Jan 27



2014 Feb 20



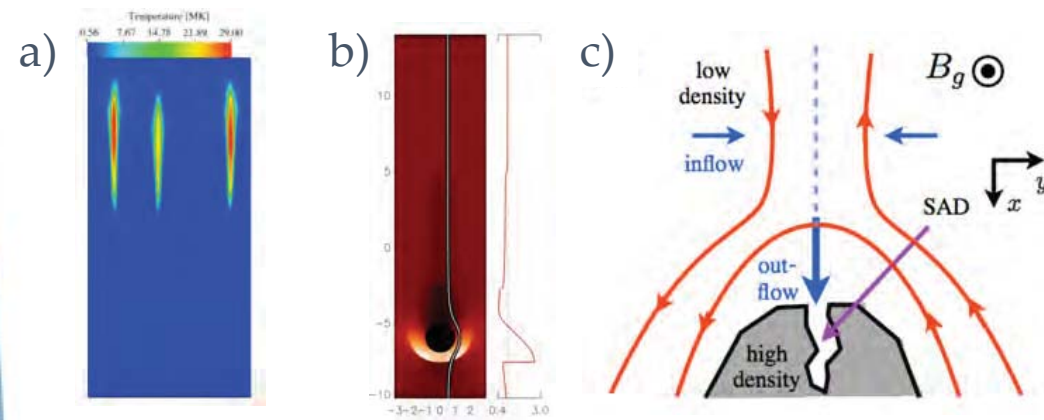


Models



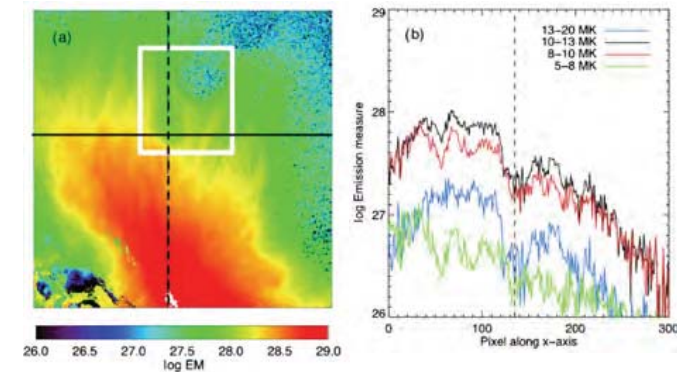
Temperature Discriminant

- ◆ a) Cecere et al 2012
 - ◆ Pressure pulse + MHD wave ($T \gg \text{fan}$)
- ◆ b) Scott et al 2013
 - ◆ Peristaltic pumping ($T \sim < \text{fan}$)
- ◆ c) Cassak et al 2013
 - ◆ ReX outflows ($T \sim < \text{fan}$)

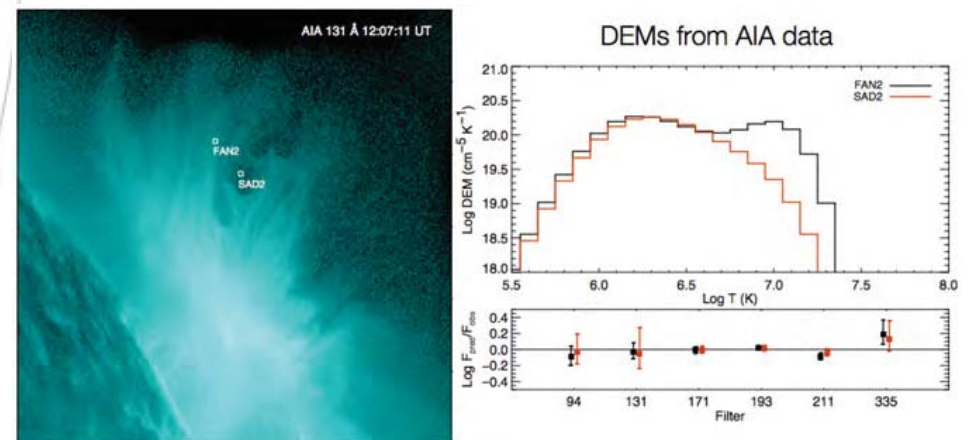


Adapted from Reeves et al Hinode 7 presentation

OBSERVED TEMPERATURE AND DENSITY ALWAYS LOWER THAN FAN

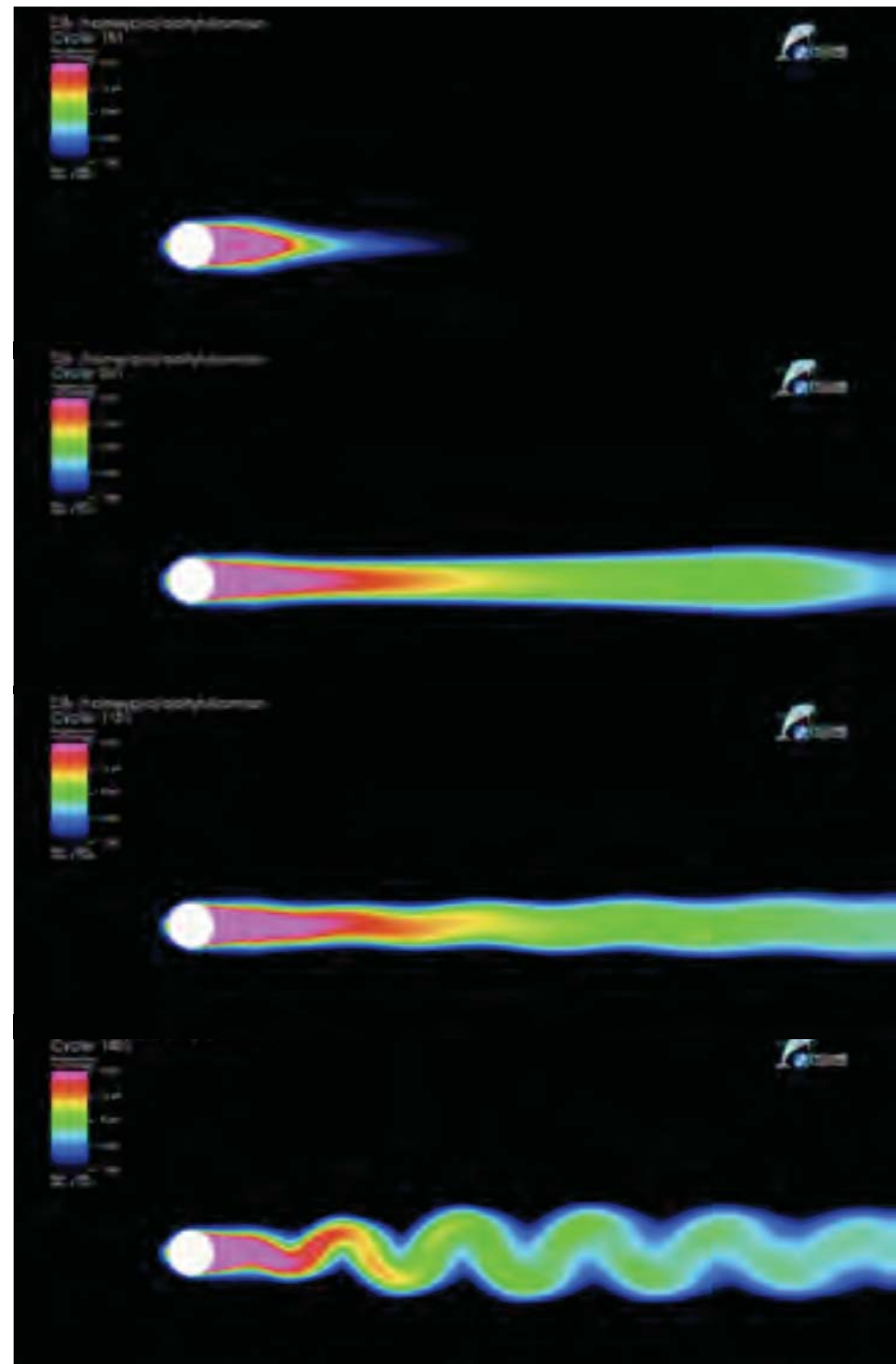


Savage, McKenzie, & Reeves 2012



Von Karman vortex street

- Perhaps something simple like some nice fluid dynamics?
-only for describing the voids and tails



Putting it all together

...or at least starting to

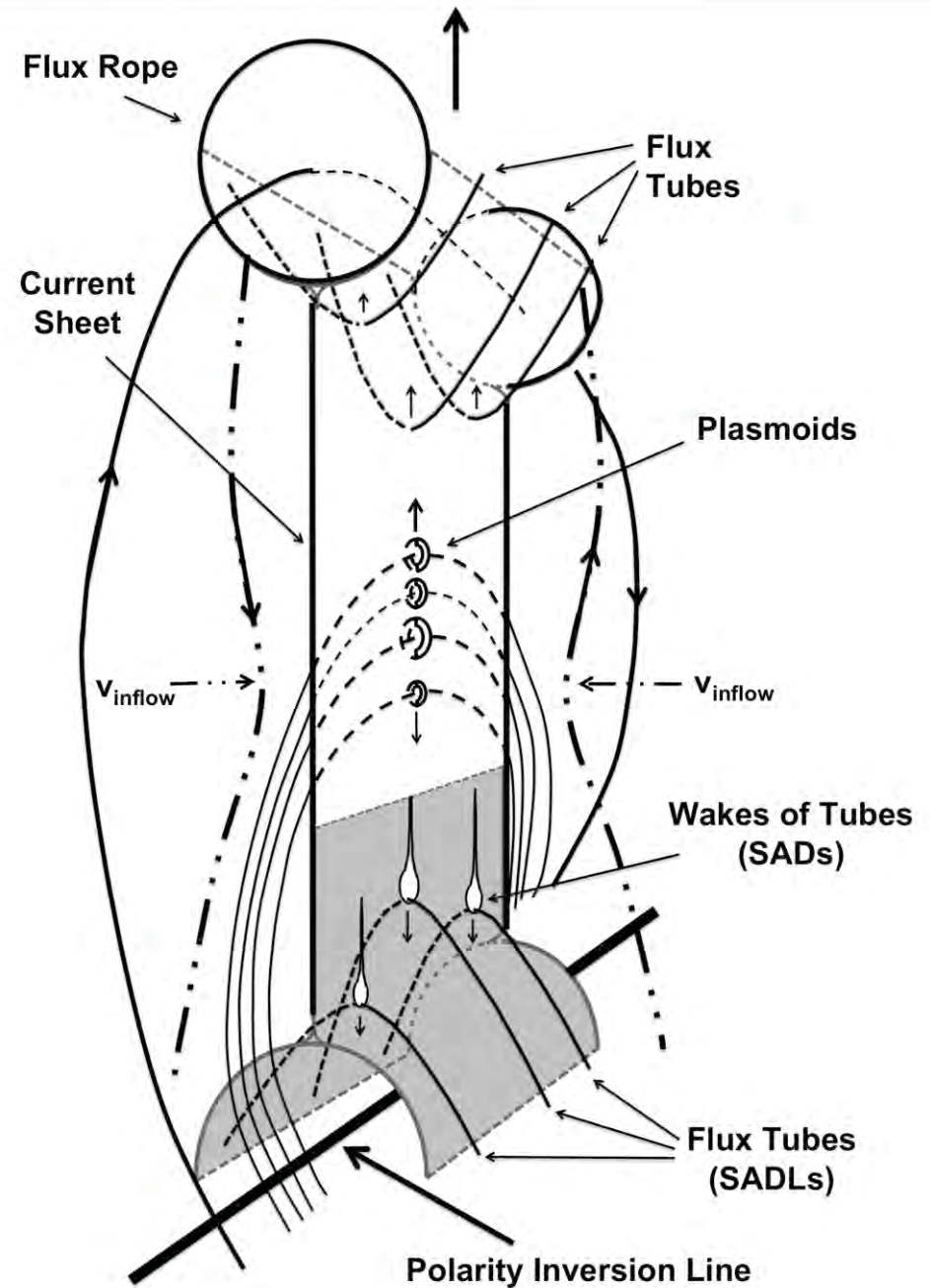


Theory

Plasmoids \neq SADs



Nishizuka & Shibata 2013, Nishida et al 2013



Savage et al 2012

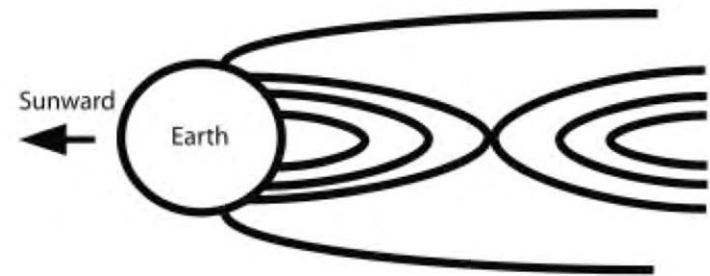
Closer to Earth

ReX counterparts in the magnetotail

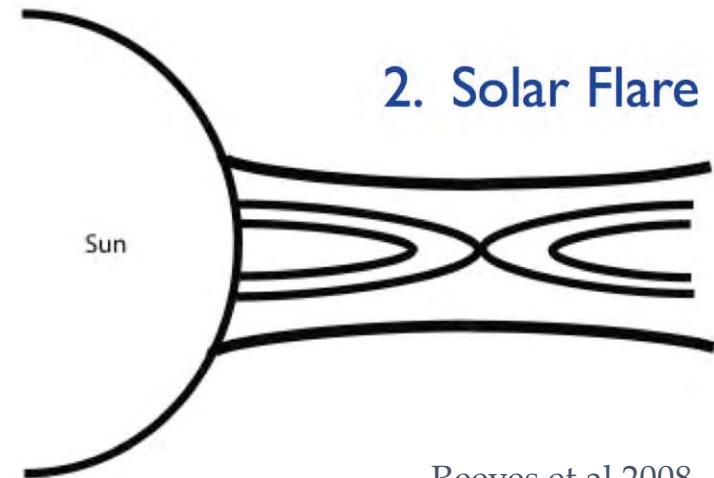


Magnetotail ReX

- Similar outputs
 - Accelerations & speeds
 - Double footpoint ribbons from electron acceleration
 - G: Aurora ribbons in dense ionosphere
 - S: Flare ribbons from chromospheric evaporation / ablation



I. Magnetotail Substorm

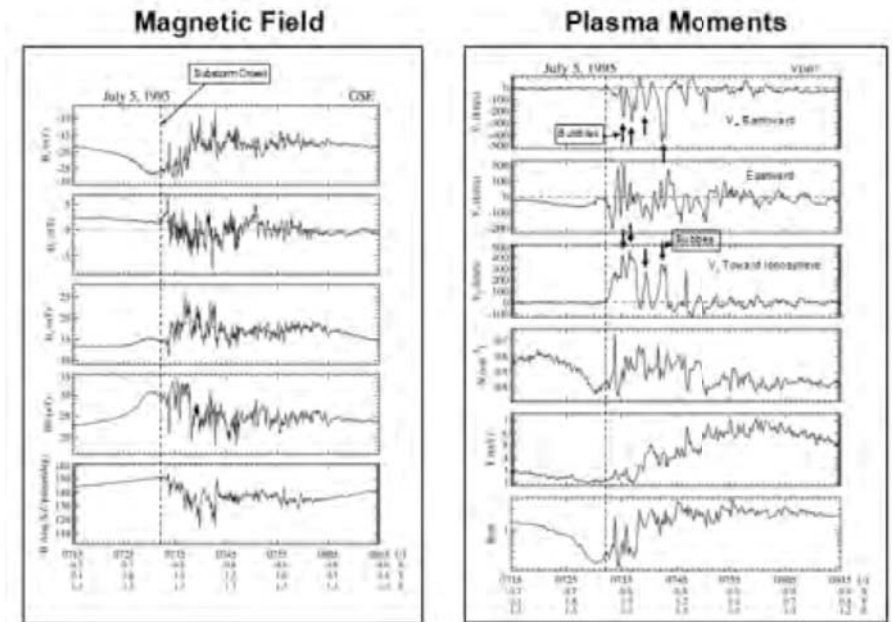


2. Solar Flare

Reeves et al 2008

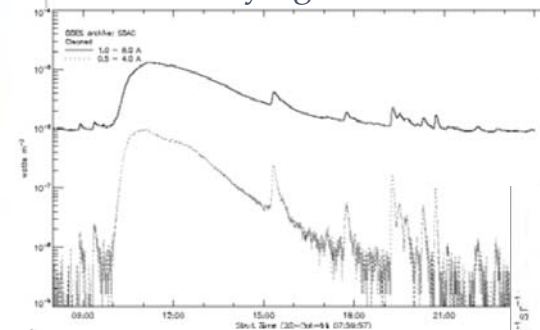
Magnetotail ReX

- ◆ Competing notation
 - ◆ G: Dipolarization
 - ◆ S: Field line shrinkage
- ◆ Complementary measurement regimes
 - ◆ G: In situ B field & plasma parameters
 - ◆ S: Global context

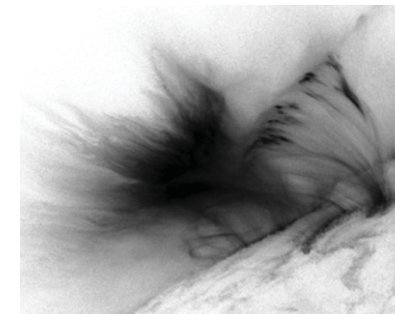
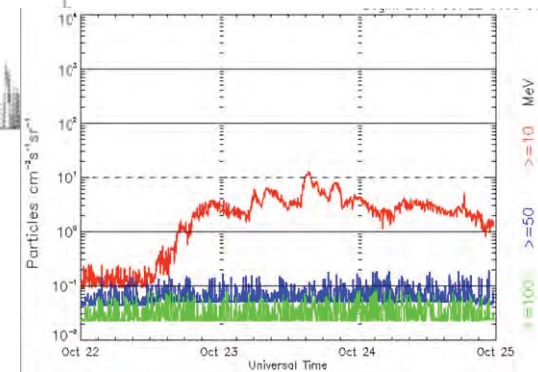


Wolf et al 2006

GOES X-ray light curve

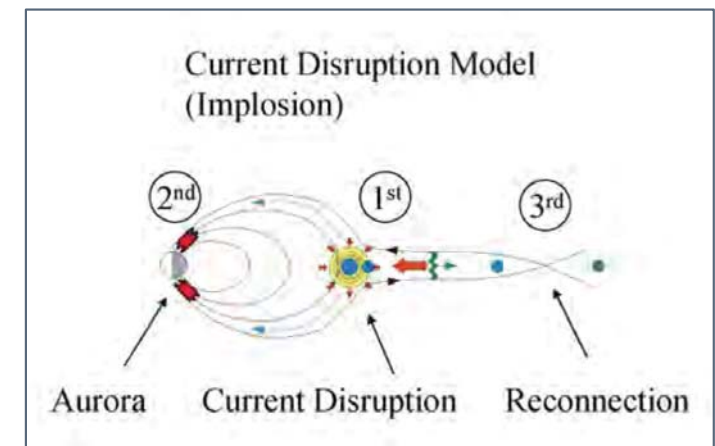
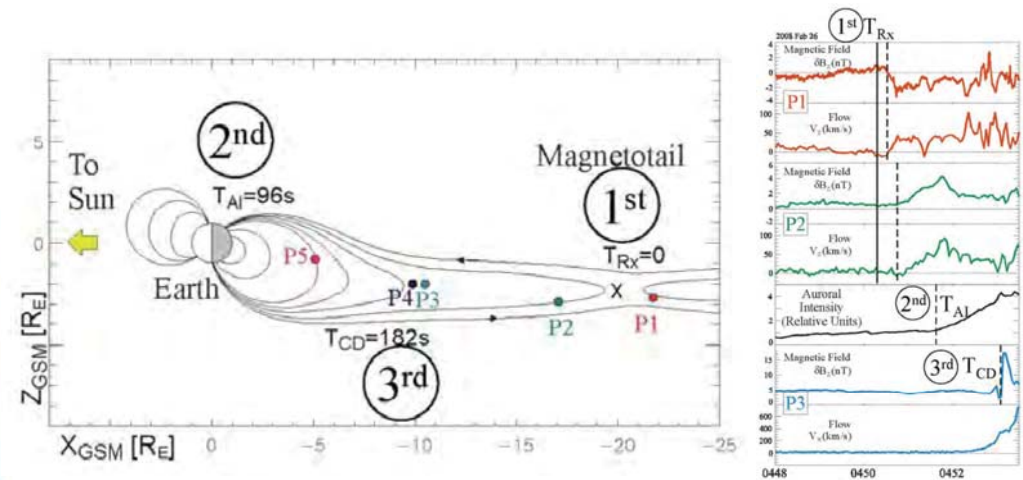


GOES Proton Flux

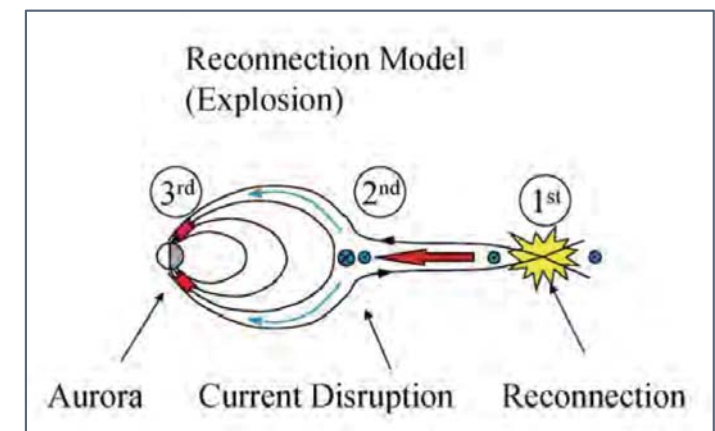


THEMIS

- 5 → 3 spacecraft
- B, E, ϕ , ...

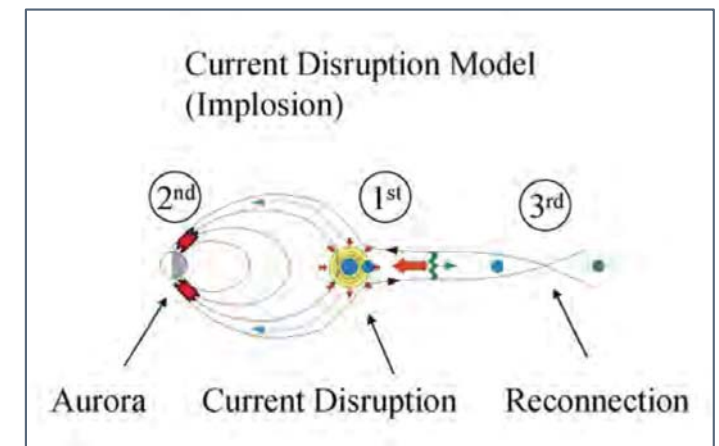
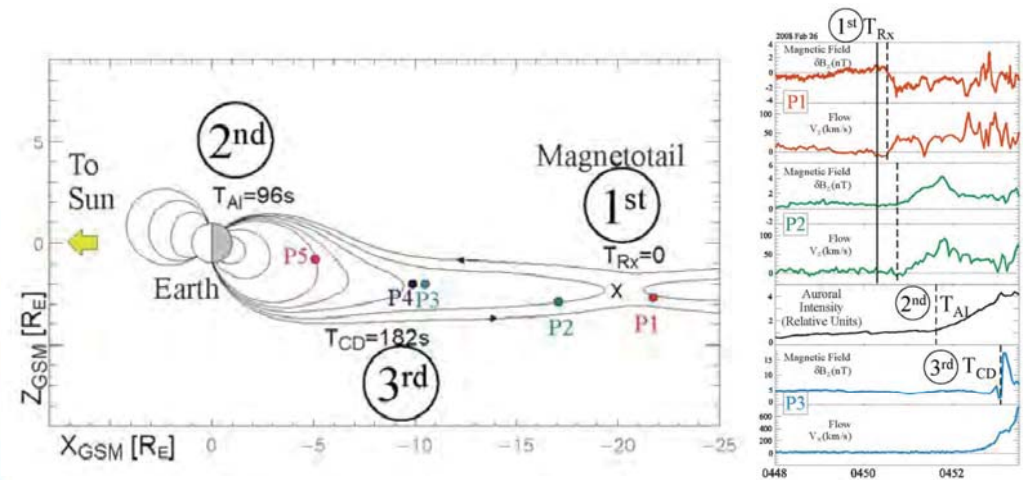


VS.

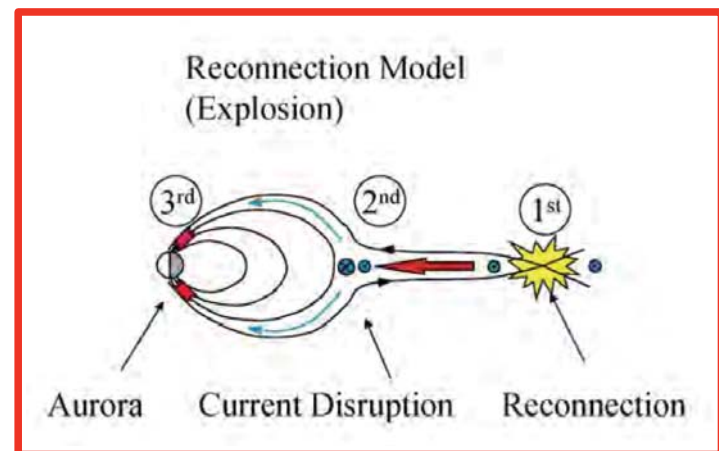


THEMIS

- 5 → 3 spacecraft
- B, E, ϕ , ...

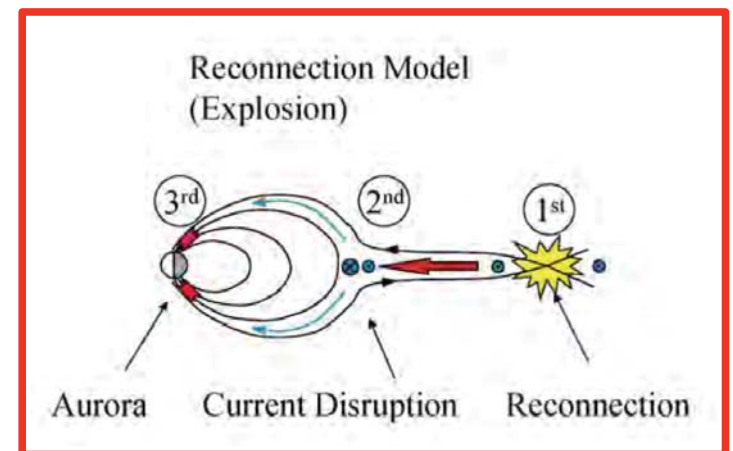


VS.



THEMIS

- 5 → 3 spacecraft
 - B, E, ϕ , ...
- SADs, Substorms, Wedgelets
 - Fast, patchy reconnection
- Similar Alfven speeds
 - But probably different regimes: resistive vs. collisionless



Summary of Outputs from Observational Properties

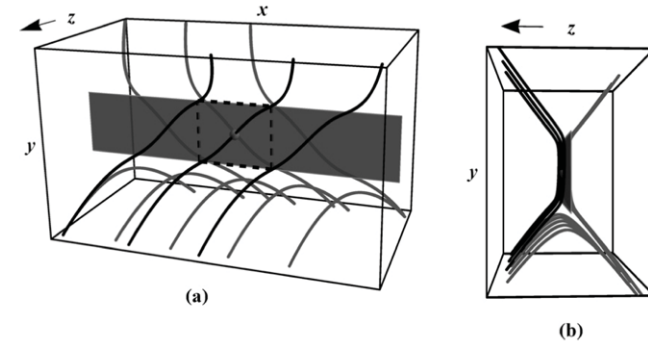
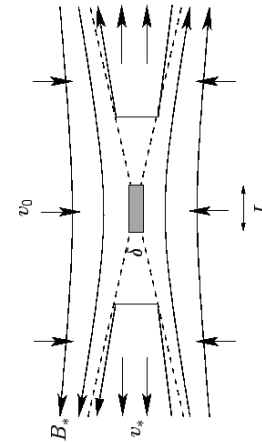
Applicable to long-duration solar flaring events



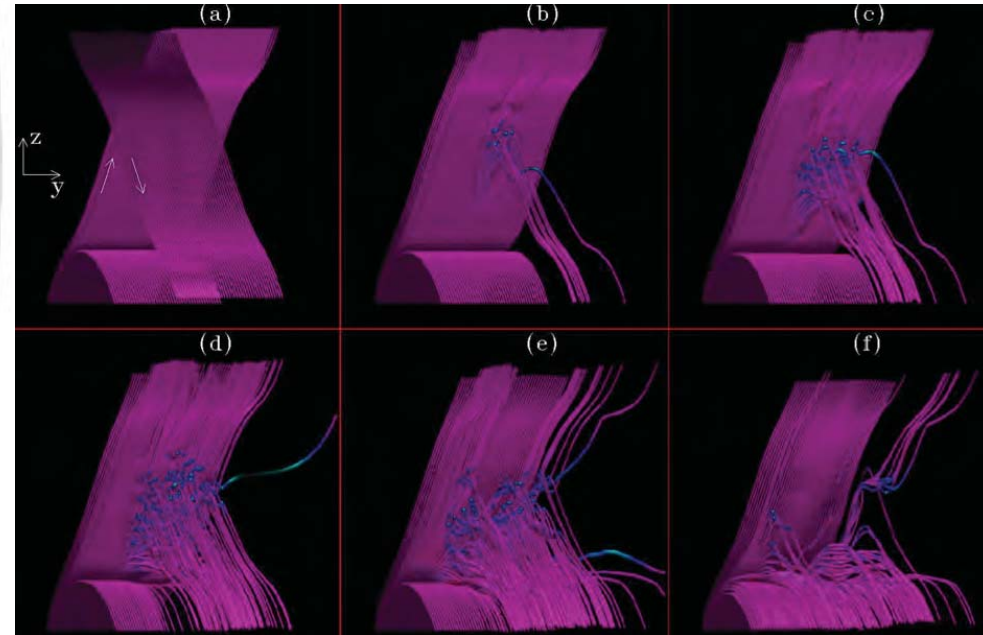
Key Extrapolated Conclusions

- ReX is fast in the presence of a guide field
 - Turbulence is also a helpful motivator
- ReX in wake of a CME is patchy & bursty (finite)

Petchek-type



Guidoni & Longcope 2010



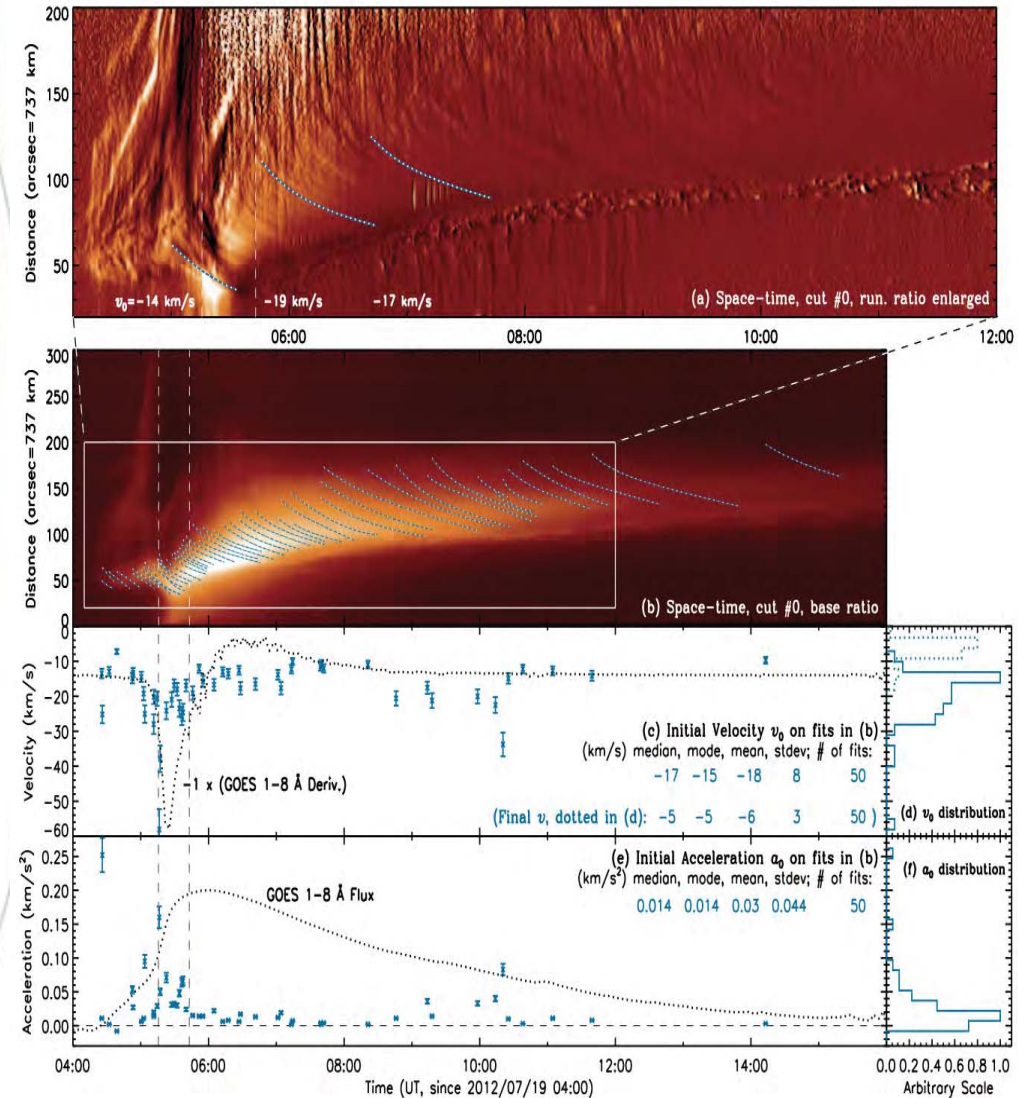
Linton & Longcope 2006

Key Extrapolated Conclusions

- Continuation of shrinking loops imparts energy into the current sheet long after the flare

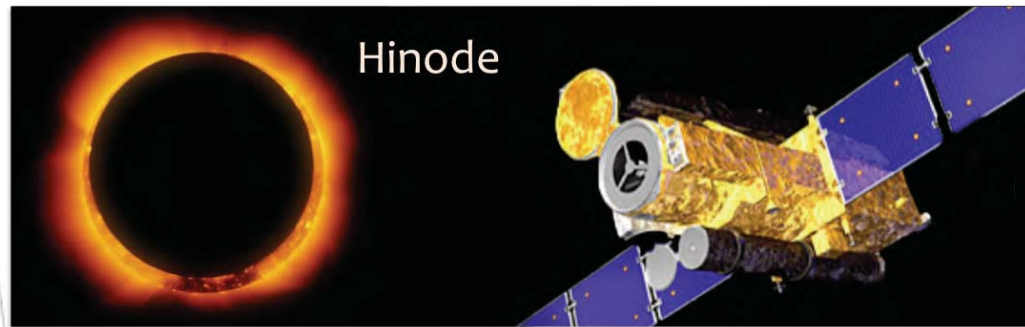
$$\Delta W = (B^2 A \Delta L / 8\pi) / \Delta t$$

per shrinking loop



Substorms + Solar flares

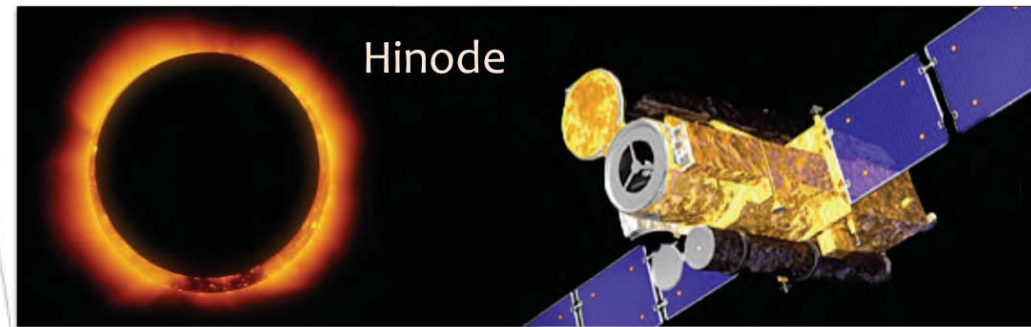
💧 Yes.



etc.

Substorms + Solar flares

- Yes.
- Potential for more complete picture of reconnection.
 - Suitable conditions?
 - How is it triggered?
 - Turbulence?
 - Etc.



etc.



Fin